HST observations of SNI987A in its early thirties

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Claes Fransson, Dennis Alp and many others ...

(2500 days)

(11,500 days)





Ring nebula and progenitor



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Circular rings inclined by ~ 45°. Kinematic age ~ 20,000 years.
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Possible formation scenarios: Interaction of BSG and RSG winds (Blondin & Lundqvist 1993), binary merger (Morris & Podsiadlowski 2009), photoevaporation of ring emitted by BSG (Smith+2013), bipolar jets (Akashi+2015).

Connected to unusual BSG progenitor [see also poster by Dennis Alp for new constraints on the progenitor].





















Apr 1999



Nov 2000



Dec 2001



Jan 2003



Nov 2003





Apr 2006



Dec 2006



May 2007



Feb 2008



Apr 2009







Jan 2011



Feb 2013



Jun 2014



May 2015



Jun 2016







Jul 2018





The rise and fall of the ring



Fading of the ring is seen also in IR (*Arendt*+2016). Flattening of soft X-ray light curve (*Frank*+2016). Re-acceleration of the shock wave in radio (*Cendes*+2018).



Asymmetries

The eastern side peaked earlier.

South-west currently three times brighter than south-east.

Consistent with highest velocities observed in spectra of the hotspots.



New spots outside the ring



- About a dozen new spots have appeared outside the ring between 2013-2018
- Strong in narrow filters so most likely line emitting
- Peak fluxes one-two orders of magnitude lower than the hotspots in the ring

Diffuse emission outside the ring





Diffuse emission outside the ring









Get new "intermediate" rings (IR) from ejecta interacting with high-latitude material.

Will look like a single ring due to overlap with the equatorial ring in the projection.



Expect deviations from this simple model. Asymmetries, more complex geometry of material connecting rings, interaction at range of velocities...

Plausible scenario: Seeing fast ejecta interact with high-latitude material with $n \sim 10^3 \text{ cm}^{-3}$

- The new spots would be clumps with transmitted shocks (~ 100 km s⁻¹)
- The rim of diffuse emission would be the reverse shock. Velocities of ~10,000 km s⁻¹ (reverse shock emission seen in spectra at these velocities)

Need spatially resolved spectra to test model. Some information from narrow filters.

Both IRs blueshifted in north and redshifted in south. Higher blueshift for southern IR.



Probes H α between -3100 and -7500 km s⁻¹. See also *France et al.* (2015)

Blueshifts higher than expected for simple model with straight lines connecting rings. Implies more curved shape.





Diffuse emission from photoionised high-latitude material at early times?

Light echoes also suggest that there is material connecting the rings (Crotts+1995, Sugerman+2005).

The ejecta



The ejecta fade as expected from radioactive decays until ~ 5000 days.

Brightening thereafter due to heating by X-rays from the ring (Larsson+2011).

The flux is now increasing at a slower rate and the western side is significantly brighter.

These trends are in agreement with the evolution of X-ray emission (*Frank*+2016).

Implications for the morphology



Powering by X-rays explains the edge-brightened morphology of the ejecta (Larsson+13, Fransson+13).

Optical emission from ejecta currently dominated by a clump in the west (Larsson+16).

X-rays do not reach the innermost ejecta. Molecules can survive!



H₂ in the ejecta (Larsson+2019)



The future of SN 1987A

The optical emission from the ring will continue to fade. Extrapolation shows that it will be gone by ~ 2035 .

Expect more emission outside the ring at all wavelengths. Gradually probing the mass-loss history of the progenitor. New insight into the formation of the rings!

The evolution of the emission from the inner ejecta is strongly coupled to the X-ray emission. We will resolve smaller spatial scales as the ejecta expand.

Detection of the compact object?!