# HYDRODYNAMIC SIMULATIONS OF SUPERNOVA REMNANTS: DUST DESTRUCTION BY THE REVERSE SHOCK

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# THE SNDUST PROJECT

### Theme 1:

Measuring dust ejecta mass from observations

### Theme 2:

Modelling of dust (destruction) in SNRs



Antonia Bevan









Roger Wesson

llse De Looze

Florian Kirchschlager Maria Niculescu- Felix Priestley Duvaz

All group members have brought posters to present their work so please do check those out! :)



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## PREVIOUS DUST DESTRUCTION MODELLING

## Typical setups

- Analytical density profiles
- Shock-ejecta (smooth/clumpy) interactions

Numerical studies<sup>1</sup> predict SNR dust survival rates anywhere between

0 and 99% but:

### Simplified physics

- Passive non-interacting tracers
- No grain-grain collisions
- Initial conditions

<sup>1</sup>e.g. Nozawa et al., 2007 (0-80%), Bianchi & Schneider, 2007 (2-20%), Nath et al., 2008 (80-99%), Silvia et al., 2010 (0-70%), Biscaro & Cherchneff, 2016 (6-98%), Boccio et al., 2016 (1-8%), and Micelotta et al., 2016 (12-16%)

# THE 'CLOUD CRUSHING' MODEL

We simulate the conditions in the SNR using the hydrodynamic code AstroBEAR (Carroll-Nellenback et al., 2013) and the 'Cloud Crushing' model (Woodward, 1976).



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### dusty fields



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### inter-cell transport



- We sort the dust grains into N grain size bins
- Each bin is added as a timedependent quantity to our HD simulations
- Modified advection routines move the dust from cell to cell
- Source terms shuffle grains between bins (dust destruction & creation)

### inter-bin transport



# **MODEL FEATURES**

## Hydrodynamic Simulations

- · Cooling
- · AMR

## Dust physics

- Carbon and/or Silicate dust
- Grain charging
- Grain size distribution tracking
- Drag (gas and plasma)
- **Sputtering** (thermal & kinetic)
- Grain-grain collisions



# <u>RESULTS</u>

# DUST SURVIVAL RATES IN CAS A

# INITIAL CONDITIONS

### Ambient Medium

$$n_{\rm am} = 1 \,{\rm cm}^3$$
  $T = 10^4 \,{\rm K}$ 

### Gas Cloud

 $n_{cl} = \chi \cdot n_{am}$   $\chi \in [100, 1000]$  $T = 10^2 \text{ K}$   $R_{cl} = 10^{16} \text{ cm}$ 

### Shock

 $v_{sh} = 1600 \text{ km/s}$ 

### Dust

$$\Delta_{gd} = 10$$





## SPUTTERING & GRAIN-GRAIN COLLISIONS



We find that the inclusion of grain-grain collisions facilitates dust destruction by turning large grains into smaller grains which are more easily sputtered.

## INITIAL DUST GRAIN SIZE DISTRIBUTIONS



Kirchschlager et al., 2019, submitted

Large grains and broad distributions result in the highest survival rates if we consider sputtering only.

## INITIAL DUST GRAIN SIZE DISTRIBUTIONS



Kirchschlager et al., 2019, submitted

Survival rates change significantly if we consider grain-grain collisions due to the **interplay between collisions** and **sputtering**.

## CLUMP-TO-AMBIENT-MEDIUM GAS DENSITY RATIO (C)



Kirchschlager et al., 2019, submitted

Depending on the initial clump-to-ambient-medium gas density ratio we find **survival rates** up to **30%** for carbon dust.



Kirchschlager et al., 2019, submitted

Depending on the initial clump-to-ambient-medium gas density ratio we find **survival rates** up to **40%** for silicate dust.



### PAPERBOATS

An external dust processing code for HD simulations.

For more information please see the **poster** & have a chat with **Florian** and keep an eye out for **Kirchschlager et al., 2019** (submitted)!

# CONCLUSIONS

### Summary

- We use hydrodynamics simulations and dust processing routines to model dust destruction in SNRs
- Grain-grain collisions facilitate dust destruction by sputtering and should not be neglected!
- Depending on grain species, grain size distribution and clump-toambient-medium gas density ratio we find survival rates up to 40%.

### Future Work

- **3D** simulations (DiRAC!)
- MHD simulations