The Velocity of Cas A's Reverse Shock and its Effects on its Ejecta

#### **Rob Fesen**

Dan Patnaude Dan Milisavljevic Kathryn Weil McKinley Brumback Cassiopeia A D = 3.4 kpcSN ~1671Age ~ 350 yrType IIb ~ $15 \text{ M}_{o}$ 

Optical/HST



**Ejecta Velocities:** 

#### -4000 to +6000 km/s

DeLaney et al. 2010

Cas A's ejecta are arranged in large "rings", marking the crosssections of ejecta bubbles encountering the reverse shock front.



Milisavljevic & Fesen

Below illustrates the basic structure of an idealized, spherically symmetric remnant: unshocked ejecta, reverse shock, shocked ejecta, contact discontinuity, shocked ambient gas, and unshocked ambient gas. This figure is taken from a simulation of uniform ejecta expanding into a uniform ambient medium.



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## Forward Shock: V ~ 5000 km/s

### Radius: 2.6 pc

Delaney & Rudnick 2003 Patnaude & Fesen 2009

# What about the velocity of the reverse shock?

# Well...that's where things get interesting.



Truelove & McKee 1999



Truelove & McKee 1999

 $V_{RS ejecta} = V_{ejecta} - V_{RS sky}$ 

Helder & Vink (2008) found for a small western limb region:

 $V_{RS sky}$  = -970 km/s

Meaning the reverse shock is moving back toward remnant center.

if  $V_{ejecta}$  = 5000 km/s, the ejecta contacts the RS with a velocity of  $V_{RS ejecta}$  = 6000 km/s



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However, elsewhere (i.e., northern and eastern regions)  $V_{RS sky} \sim +2000 \text{ km/s}$ ; meaning it moves outward wrt sky frame

Larger transverse proper motions have been reported across central and western regions for Cas A's nonthermal filaments.



Vector length is proportional to the actual shifting value. Blue and red show outward and inward motions, respectively.

#### Inward transverse velocities: -2100 to -3800 km/s (!)

Sato+ 2018

Consequently, they claim:

 $V_{RS ejecta} = 5100 - 8700 \text{ km/s}$ 

which is nearly 3000 km/s faster than the ISM/CSM sees the forward shock.

Sato et al. discount possible inward motions as reverse shocks (in the observer's reference frame) for an evolutional phase that describes a 350 yr old Cassiopeia A...because the reverse shock should be expanding outward.

Instead, they proposed the inward shock was a "reflection shock" produced by the forward shock interacting with a density jump in the ambient medium.



Kilpatrick+ 2014 But see poster Zhou+ S10.21

# What's Cas A's reverse shock velocity when measured in the optical?

Well... its not so simple

There has been significant changes in Cas A's optical appearance over the last 60 yrs.

Due to the advancement of the reverse shock into the expanding ejecta cloud.



Patnaude & Fesen 2012

#### Changes in Cas A with time























Ejecta Velocities

0.34" - 0.37"/yr 5400 - 5900 km/s

0.32" - 0.34"/yr 5100 - 5400 km/s

> 0.30" - 0.32"/yr 4900 - 5100 km/s

# RS Velocities

-0.05"/yr \_\_\_\_\_ -800 km/s

-0.02"/yr\_\_\_\_ -300 km/s

> 0.08"/yr\_\_\_\_ 1300 km/s

> > 0.12"/yr 1900 km/s











RS Velocity -0.04"/yr to -0.06"/yr -650 to -1000 km/s

> Ejecta Velocities 0.25"/yr to 0.35"/yr 4000 to 5600 km/s




Knot - Location of Reverse Shock





<u>Knot</u> <u>Velocity:</u> 0.32"/yr = 5150 km/s



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<u>RS Velocity</u> (Base) 0.0"/yr = 0 km/s



<u>Knot</u> <u>Velocity:</u> 0.32"/yr = 5150 km/s

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So: V<sub>RS ejecta =</sub> 5150 km/s (projected velocity)













2"\*

Reverse Shock Velocity 0.11"/yr ==> 1800 km/s

#### 2004.93 ----

Ejecta Velocity 0.25"/yr ==> 4000 km/s

## 2000.06

There are strong dynamical effects of Cas A's reverse shock on the structure of its ejecta.



5"

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# Knot 2 Knot 3

# Knot 1

15"

# Knot 4





## Take Aways

1) The velocity of Cas A's reverse shock seems to depend on the density of ejecta knots it is moving around and through. -- ejecta dense regions  $V_{RS ejecta} = 1000 - 3000 \text{ km/s}$ -- lower density regions  $V_{RS ejecta} = 5000 - 6000 \text{ km/s}$ 

RS radius wrt COE: 1.4 pc – 1.8 pc

Why is the reverse shock velocity so high? Unclear. Interaction of the forward shock with dense CSM clouds, along with varying ejecta density profiles (n=7 to ?) may play a role in increasing Cas A's reverse shock velocity.

2) Cas A's reverse shock front strips the outer layers of small dense ejecta knots creating long, contrail-like streams (~0.008 pc long and slower (500 – 1000 km/s) of higher ionized material.

### Slit 3

Blueshift

#### [O III] 4959 5007 •

#### Knot 3v = -1800 km/s

#### Redshift

# $\begin{array}{ccc} & 4959 & 5007 \\ \text{Knot 4} & \longrightarrow & & & \\ v = +4800 \text{ km/s} & \text{[O III]} \end{array}$

km/s

1500 km/s 25 Ang






2000.06

5"

2019.07

5"



.1951



Figure 6. Extracted spectra for forward shock and interior nonthermal filaments.



D. Milisavljevic

Surface of inner caviety of unshocked Ejecta









